Classical Thermodynamics

Ideal gases

- Assumptions
 - o Identical particles in random motion.
 - Small solid spheres occupy negligible volume.
 - Elastic collisions.
 - No long range forces only forces during collisions \rightarrow only energy is KE, $\frac{1}{2}mv^2$ per particle.
- Results
 - Pressure is given by $p = \frac{1}{3} nm \langle v^2 \rangle$.
 - o Flux is given by $J = \frac{1}{4}n \langle v \rangle$.
- The Maxwell-Boltzmann Distribution
 - o $P(v) = \left(\frac{m}{2\pi kT}\right)^{3/2} 4\pi v^2 \exp\left(-mv^2/2kT\right)$
 - o This gives

- Energy...
 - $\circ \quad U = \frac{3}{2}RT = \frac{3}{2}N_A kT$

Definitions

- A system is in Thermal Equilibrium when all its macroscopic observables have ceased to change with time.
- A function of state is any physics quantity that has a well-defined value for each equilibrium state of the system. They are represented by exact differentials.
- Functions of state can be either
 - \circ Extensive (proportional to system size) eg: volume.
 - \circ Intensive (independent of system size) eg: pressure.
 - Intensive and extensive variables form **conjugate pairs**, the product of which is **energy**.

• The heat capacity is the amount of heat we need to supply to raise the temperature by dT. This can be measured at constant volume or at constant pressure, so

$$C_{V} = \left(\frac{\mathrm{d}\,Q}{\mathrm{d}\,T}\right)_{V} \qquad \qquad C_{p} = \left(\frac{\mathrm{d}\,Q}{\mathrm{d}\,T}\right)_{p}$$

• For an **ideal gas**

$$pV = nRT$$

• Stirling's Approximation

$$\ln N \,! \approx N \ln N - N$$

The First Law

• The First Law of Thermodynamics states that

Energy is conserved, and heat and work

are both forms of energy

$$\Delta U = \Delta Q + \Delta W$$

- Convention:
- ΔQ and ΔW are POSITIVE when energy is given TO the system
- For a differential change

$$\mathrm{d}\,U = \mathrm{d}\,Q + \mathrm{d}\,W$$

• The work done compressing a gas is given by

 $\frac{\mathrm{d} W}{\mathrm{d} W} = -p \mathrm{d} V$ And so **for a gas**, the first law can be written

$$\mathrm{d}\,U = \mathrm{d}\,Q - p\,\mathrm{d}\,V$$

Heat Capacities

• In general, the **internal energy** will be a function of **temperature** and **volume**, so

$$\mathrm{d} U = \left(\frac{\partial U}{\partial T}\right)_V \mathrm{d} T + \left(\frac{\partial U}{\partial V}\right)_T \mathrm{d} V$$

• Using the form of the **first law** for a **gas**, above, we can write

$$\mathrm{d} Q = \mathrm{d} \, U + p \, \mathrm{d} \, V$$

$$\mathbf{d}Q = \left(\frac{\partial U}{\partial T}\right)_{V} \mathbf{d}T + \left[\left(\frac{\partial U}{\partial V}\right)_{T} + p\right] \mathbf{d}V$$

And dividing by dT:

$$\frac{\mathrm{d}Q}{\mathrm{d}T} = \left(\frac{\partial U}{\partial T}\right)_{V} + \left[\left(\frac{\partial U}{\partial V}\right)_{T} + p\right] \frac{\mathrm{d}V}{\mathrm{d}T}$$

• By taking the equation above at constant volume and constant pressure (only the dV/dT term will matter), we obtain

$$\begin{split} C_{V} &= \left(\frac{\partial U}{\partial T}\right)_{V} \\ C_{p} &= C_{V} + \left[\left(\frac{\partial U}{\partial V}\right)_{T} + p\right] \left(\frac{\partial V}{\partial T}\right)_{p} \end{split}$$

• For an ideal gas:

$$C_{p} = C_{V} + R$$

• We define

$$\gamma = \frac{C_p}{C_V}$$

Reversibility

- A change is **reversible** if an **infinitesimal change in external conditions** would **reverse the direction of the change**.
- Reversible changes are typically **very slow** and **quasi-static**.
- They are also **frictionless** no **viscosity**, **turbulence**, etc...

Isothermal expansions

• When an expansion is **isothermal**, **the temperature of the system does not change**. Therefore, the **internal energy** of the system does not change, and

$$\frac{\mathrm{d}}{\mathrm{d}}W = -\frac{\mathrm{d}}{\mathrm{d}}Q$$

• Therefore, when a gas is **expanded isothermally** from V_1 to V_2 at a temperature T, the **heat absorbed by the gas** is given by

$$\Delta Q = \int d Q$$

= $-\int d W$
= $\int_{V_1}^{V_2} p d V$
= $R T \int_{V_1}^{V_2} \frac{1}{V} d V$

$$\Delta Q = RT \ln \frac{V_2}{V_1}$$

• An adiabatic process is both adiathermal (no flow of heat) and reversible, so

$$dQ = 0$$

And

$$\mathrm{d}\,U = \mathrm{d}\,W$$

However, for an ideal gas

$$\mathrm{d}\, U = C_V \mathrm{d}\, T$$

Therefore

$$\begin{split} C_V \mathrm{d}\, T &= \mathrm{d}\, W \\ C_V \mathrm{d}\, T &= -p \,\mathrm{d}\, V \\ C_V \mathrm{d}\, T &= -\frac{R\,T}{V} \,\mathrm{d}\, V \\ \frac{\mathrm{d}\, T}{T} &= -\frac{R}{C_V} \frac{\mathrm{d}\, V}{V} \\ \frac{\mathrm{d}\, T}{T} &= \frac{0 \,\mathrm{Divious from}}{\left(1-\gamma\right)} \,\frac{\mathrm{d}\, V}{V} \\ \frac{\mathrm{d}\, T}{V} &= \overline{\left(1-\gamma\right)} \,\frac{\mathrm{d}\, V}{V} \end{split}$$

Other versions can easily be generated using $pV \propto T$.

The Second Law

• The Second Law of Thermodynamics can be stated in two different ways

The Clausius Formulation

No process is possible whose sole result is the transfer of heat from a hotter to a cooler

body.

The Kelvin Formulation

No process is possible whose sole result is the complete conversion of heat into work.

- The **equivalence** of these two formulations can be shown in two steps
 - $\circ \quad \text{Violating Kelvin} \rightarrow \text{Violating Clausius}$

Consider a "Kelvin-violator" E connected to a Carnot engine:



We then have

$$Q'_h = W$$
$$Q_h = W + Q_l$$

But the total heat dumped at T_h is given by

$$Q_h - Q'_h = W + Q_l - W = Q_l$$

Thus, the **net effect** is to **transfer heat from** T_l to T_h – this

violates Clausius' Formulation.

◦ Violating Clausius → Violating Kelvin

Consider a "Clausius violator" *E* running next to a Carnot engine:



Overall, no heat is being dumped at T_l , but a heat $Q_h - Q_l$ is drawn from T_h . Furthermore, by the First Law

$$W = Q_h - Q_l$$

Which means that the **net effect** of this engine is to convert heat $Q_h - Q_l$ into work which Violates Kelvin's Formulation.

Engines

- Engines are systems operating cyclic processes that convert heat into work.
- The Carnot Engine is based on the Carnot Cycle:
 - o Isothermal expansion heat Q_h absorbed.

- \circ Adiabatic expansion no heat flow.
- o Isothermal compression heat Q_l given out.
- Adiabatic compression no heat flow

A schematic is as follows:



Where S is some quantity proportional to pV^{γ} , constant during an adiabatic change.

• By the **First Law**, the **work output** must be

$$W = \oint p \,\mathrm{d}\, V = Q_h - Q_l$$

• The **efficiency** of an engine is defined as

$$\eta = \frac{\text{What we want}}{\text{What we gave for it}} = \frac{\text{Work done by engine}}{\text{Heat put in}} = \frac{Q_h - Q_l}{Q_h} = 1 - \frac{Q_l}{Q_h}$$

• For a **Carnot Cycle**, we can show, using the formulae derived above for isotherms and adiabats, that

$$\frac{Q_{\scriptscriptstyle h}}{Q_{\scriptscriptstyle l}} = \frac{T_{\scriptscriptstyle h}}{T_{\scriptscriptstyle l}} \Rightarrow \eta = 1 - \frac{T_{\scriptscriptstyle l}}{T_{\scriptscriptstyle h}}$$

• Note, crucially, that this implies that the quantity Q/T is <u>always</u> the same in a reversible Carnot Cycle, so, for a **Carnot Cycle**

$$\oint \frac{\mathrm{d}Q}{T} = 0$$

Furthermore, for any other sort of engine, $\eta_{Engine} \leq \eta_{Carnot}$, so

$$\begin{split} 1 - \frac{Q_l}{Q_h} &\leq 1 - \frac{T_l}{T_h} \Rightarrow \frac{Q_h}{T_h} \leq \frac{Q_l}{T_l} \\ & \boxed{\oint \frac{\mathrm{d} Q}{T} \leq 0} \end{split}$$

We can show that this is true for *any* cycle – **Clausius' Theorem** (see below).

- An engine can be run in reverse in two ways:
 - o Refrigerator

- In this case, the desired aim is the removal of heat from a cold reservoir, and this is done by the provision of work
- Therefore, the **efficiency** is given by

$$\eta = \frac{Q_l}{W}$$

- High efficiency can be achieved if $T_2 \approx T_1$.
- o Heat pump
 - In this case, the desired aim is the dumping of heat at a hot reservoir, and this is done by the porivision of work.
 - Therefore, the efficiency is given by

$$\eta = \frac{Q_h}{W}$$

• This is always greater than 1, because **all work** is converted to heat (eg: like in an electric heater), but **more** is also added, from the **cold reservoir**!

Carnot's Theorem

• Carnot's Theorem states that

<u>Carnot's Theorem</u> Of all the heat engines working between two given temperatures, none is more efficient than a Carnot engine

Proof



Consider an engine E that is **more efficient** than a Carnot Engine (ie: $\eta_E > \eta_{Carnot}$), whose **produced work** is used to drive a **Carnot** engine **in reverse**. We then have

$$\begin{split} \eta_E &> \eta_{Carnot} \\ \frac{W}{Q_h'} &> \frac{W}{Q_h} \\ Q_h &> Q_h' \\ Q_h &- Q_h' &> 0 \end{split}$$

The First Law further implies that

$$W = Q_h^\prime - Q_k^\prime = Q_h - Q_l$$

And therefore

$$Q_h - Q_h' = Q_l - Q_l' > 0$$

This means that the **net effect** of these engines is to **transfer energy**

<u>from</u> T_l to T_h – this violates the Clausius Formulation.

• A corollary is that

Corollary to Carnot's Theorem

All reversible engines have the same

efficiency as a Carnot engine, η_{Carnot}

Proof



We now imagine a reversible engine R, with $\eta_R < \eta_{Carnot}$ and drive it backwards, using the work from a Carnot Engine. The overall result will be to transfer heat from T_l to T_h , in violation of Clausius' Formulation. Thus, since the efficiency of any engine has to be both \leq and \geq that of a Carnot engine, it must be equal.

Clausius' Theorem

- Consider **any cycle** that
 - Absorbs heats dQ_i at various points i which have temperature T_i .
 - **Releases**, overall, work ΔW , which, by the **First Law**, must be

$$\Delta W = \sum_{\text{cycle}} \mathbf{d} Q_i$$

• Further imagine that each of these heats, dQ_i , are being provided by **Carnot Engines**, operating from a **single reservoir** at T_o and product work dW_i each time. From the discussions above, we know that

$$\begin{split} \frac{\mathrm{d}Q_i}{T_i} &= \frac{\mathrm{d}Q_i - \mathrm{d}W_i}{T} \\ \mathrm{d}W_i &= \mathrm{d}Q_i \bigg(\frac{T}{T_i} - 1\bigg) \end{split}$$

- [Note that no assumptions has been made about the sign of dQ_i , so heat could have been released as well].
- In any case, it looks, at the moment, like all the system is doing is **absorbing heat** and **producing work**. This violates **Kelvin's Statement**. To remedy to this, we must have

total work produced per cycle =
$$\Delta W + \sum_{\text{cycle}} \frac{d}{W_i} \leq 0$$

 $T \sum_{\text{cycle}} \frac{dQ}{T_i} \leq 0$

Since T > 0, we must therefore have

$$\oint \frac{\mathrm{d} Q}{T_i} \leq 0$$

 If the cycle was reversible, we could have run it backwards (with dQ'_i = -dQ_i), and obtained the result with the inequality sign flipped. So, Clausius' Theorem is that

$$\oint \frac{\mathrm{d} Q}{T_i} \leq 0$$

For any closed cycle, with equality when the cycle is reversible.

Entropy

• We saw that for a reversible cycle, $\int_{cycle} dQ_{rev} / T = 0$, which means that the integral $\int_{A}^{B} dQ_{rev} / T = 0$ is path independent. As such, the quantity dQ_{rev} / T is an exact differential, and we can write a new state function, entropy, S, which we define by

$$\mathrm{d}S = \frac{\mathrm{d}Q_{rev}}{T}$$

• Now, consider going from $A \rightarrow B$ reversibly, and then returning from $B \rightarrow A$ irreversibly. We then have

$$\begin{split} & \oint_{A \to B \neq A} \frac{\mathrm{d}Q}{T} \leq 0 \\ & \oint_{A \to B} \frac{\mathrm{d}Q_{rev}}{T} + \oint_{B \neq A} \frac{\mathrm{d}Q}{T} \leq 0 \\ & \oint_{B \neq A} \frac{\mathrm{d}Q}{T} \leq \oint_{B \neq A} \frac{\mathrm{d}Q}{T} \\ & \oint_{B \neq A} \frac{\mathrm{d}Q}{T} \leq dS \end{split}$$

This is true however close A and B get to each other, so

dS >	$dQ_{\rm irreversible}$
ub ≥	Т

• However, for a thermally isolated system, dQ = 0 for any process. Therefore, assuming that the universe is thermally isolated:

$$\mathrm{d}S_{\mathrm{universe}} \geq 0$$

• We can **re-write the First Law** as follows:

$$\mathrm{d}\,U = \mathrm{d}\,Q + \mathrm{d}\,W$$

But for a reversible change in a gas, we can write

$$\mathrm{d}\,U = T\mathrm{d}S - p\mathrm{d}\,V$$

However, all the quantities in this equation are **equations of state** – therefore, this holds for **any change**, reversible or irreversible.

This implies that U changes when either S or V change – in other words, S and V are the natural variables of U, and they are both extensive – they depend on the size of the system. T and p, by contrast, are variables that are intensive, and that tell us how each of these extensive variables affect U. In fact:

Pot-Pourri of facts

- It is interesting to note that for an **adiabatic** (reversible adiathermic) change, $dQ_{rev} = 0$, and so dS = 0.
- Body of heat capacity C, from $T_1 \rightarrow T_2$:

$$\Delta S = \int_{T_1}^{T_2} \frac{\mathrm{d}Q}{T} = \int_{T_1}^{T_2} \frac{C}{T} \,\mathrm{d}T = C \ln \left(\frac{T_2}{T_1}\right)$$

• Reservoir at temperature T absorbing heat q

$$\Delta S = \frac{q}{T}$$

[Note: q is always reversible for a reservoir, because the heat content is so large that any small change can reverse the direction of flow].

• For a **ideal gas**, we can write

$$dU = TdS - pdV$$

$$\underbrace{C_V dT}_{\text{Only for gas}} = TdS - pdV$$

$$dS = C_V \frac{dT}{T} + p\frac{dV}{T}$$

$$dS = C_V \frac{dT}{T} + \underbrace{R\frac{dV}{V}}_{V}$$

$$\overline{S = C_V \ln T + R \ln V + S_0}$$

This gives the **Entropy of an Ideal Gas**, from which we can work out the **Entropy of mixing**.

This expression, however, has an **intrinsic problem** – S is **extensive**, but V is also extensive in the expression above (T is not).

Joule Expansion from V → nV - can be worked out either using the expression for S for a gas, or using the fact that S is a state function, and that ΔS will therefore be the same than for an isothermal expansion from V to nV. It turns out it's

$$\Delta S = R \ln n$$

- This is a generally very useful principle we can use a reversible change between the same two steps to find the change in entropy in an irreversible change.
- The latent heat is the amount of heat needed to convert unit mass/mole from one phase to another;

$$L = \Delta Q = T(S_2 - S_1)$$

• We can also express **heat capacities** in terms of **entropy**

$$C_{V} = \left(\frac{\mathrm{d}Q}{\mathrm{d}T}\right)_{V} = T\left(\frac{\partial S}{\partial T}\right)_{V}$$
$$C_{p} = \left(\frac{\mathrm{d}Q}{\mathrm{d}T}\right)_{p} = T\left(\frac{\partial S}{\partial T}\right)_{p}$$

• The work that can be gotten out of a system through some change is given by (note the change in signs due to the fact we're extracting energy)

$$\frac{\mathrm{d}}{\mathrm{d}}W = \frac{\mathrm{d}}{\mathrm{d}}Q - \mathrm{d}\,U > T\mathrm{d}S - \mathrm{d}\,U$$

Clearly, the maximum work can be extracted when the equality holds – ie: when the process is reversible. Increase in entropy (ie: irreversibility) corresponds to conversion of energy into a less useful form than work.

Statistics

• We note that dU = TdS - pdV implies that

$$\begin{split} T &= \left(\frac{\partial U}{\partial S}\right)_{\!\!\!V} \\ \frac{1}{T} &= \left(\frac{\partial S}{\partial U}\right)_{\!\!\!V} \end{split}$$

If we compare this with the definition of temperature, below

$$\frac{1}{k_B T} = \frac{\mathrm{d}\ln\Omega}{\mathrm{d}E}$$

We find good justification to say

$$S = k_B \ln \Omega$$

This is an expression for a system that is in a **particular macrostate** (ie: fixed energy) with an associated number of **microstates**. This is the **microcanonical ensemble**. However...

• ... we can also consider the **canonical ensemble**, a system which can be in **any number of macrostates of energy** ε_i (ie: any energy – for example, by allowing it to **exchange** energy with a **reservoir**) each of which have n_i indistinguishable microstates associated with them, where the **total number of microstates** is, necessarily, $N = \sum n_i$.

The number of ways of rearranging the system is

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$$\Omega = \frac{N!}{\prod_i n_i!}$$

And this means that the **entropy** is given by

$$\begin{split} S &= k \left[\ln N! - \sum_{i} n_{i}! \right] \\ &\approx k \left[N \ln N \stackrel{\text{Insignificant}}{\longrightarrow} - \sum_{i} n_{i} \ln n_{i} \stackrel{\text{Insignificant}}{\longrightarrow} \right] \\ &\approx k \left[N \ln N - \sum_{i} n_{i} \ln n_{i} \right] \\ &= k \left[\sum_{i} n_{i} \ln N - \sum_{i} n_{i} \ln n_{i} \right] \\ &= k \left[\sum_{i} n_{i} (\ln N - \ln n_{i}) \right] \\ &= -kM \left[\sum_{i} \frac{n_{i}}{N} \ln \left(\frac{n_{i}}{N} \right) \right] \\ &= -kM \sum_{i} P_{i} \ln P_{i} \end{split}$$

Where P_i is the **probability** of finding the system in **macrostate** *i*. And therefore, the **entropy** in any **one microstate** (ie: **actual state of existence**) of the system is

$$S = -k \sum_i P_i \ln P_i$$

• We can use this to **elegantly** derive the **Boltzmann Distribution**. What we effectively want to do is to **maximise** S/k (for tidiness) subject to

$$\sum_{i} P_{i} = 1 \qquad \text{and} \qquad \sum_{i} P_{i} \varepsilon_{i} = U$$

So, consider:

$$L = \sum_{i} -P_{i} \ln P_{i} - \alpha \left(P_{i} - 1\right) - \beta (P_{i}\varepsilon_{i} - U)$$

Choose one of the probabilities P_i :

$$\begin{split} \hline \frac{\partial L}{\partial P_j} &= 0 \\ \frac{\partial L}{\partial P_j} &= - \big[\ln P_j + 1 \big] - \alpha - \beta \varepsilon_j = 0 \\ \Rightarrow P_j &= \frac{e^{-\beta \varepsilon_j}}{Z} \end{split}$$

To find β , we note that

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$$\begin{split} S &= -k \sum_{i} \frac{e^{-\beta \varepsilon_{i}}}{Z} \ln \frac{e^{-\beta \varepsilon_{i}}}{Z} \\ &= -k \left[\sum_{i} \frac{e^{-\beta \varepsilon_{i}}}{Z} \ln e^{-\beta \varepsilon_{i}} - k \ln Z \sum_{i} \frac{e^{-\beta \varepsilon_{i}}}{Z} \right] \\ &= k \left[\beta \sum_{i} P_{i} \varepsilon_{i} + k \ln Z \right] \\ &= k \beta U + k^{2} \ln Z \end{split}$$

Therefore:

$$k\beta = \left(\frac{\partial S}{\partial U}\right) = \frac{1}{T}$$
$$\beta = \frac{1}{kT}$$

As, indeed, expected.

Analytical Thermodynamcis

Thermodynamics Potentials

- We can make a number of other functions of state by adding to U various other combinations of the functions of state p, V, T and S such that the resulting quantity has energy dimensions.
- Most, however, are **not very useful**. A number, however, **are**.
- INTERNAL ENERGY dU = TdS pdV

For processes at a constant volume (isochoric), we have

$$\mathrm{d}\,U = T\mathrm{d}S = \mathrm{d}Q_{\mathrm{rev}} = C_V\mathrm{d}\,T$$

As such, if a process involves going from T_1 to T_2 at constant V:

$$\Delta U = \int_{T_1}^{T_2} C_V \,\mathrm{d}\, T$$

- **Enthalpy** $H = U + PV \Rightarrow dH = TdS + Vdp$
 - o The natural variables are now entropy and pressure.
 - o For isobaric processes (at constant pressure)

$$\mathrm{d}H = T\mathrm{d}S = \mathrm{d}Q_{\mathrm{rev}} = C_p\mathrm{d}T$$

As such, if a process involves going from T_1 to T_2 at constant p:

$$\Delta H = \int_{T_1}^{T_2} C_p \,\mathrm{d}\, T$$

o Also

$$T = \left(\frac{\partial H}{\partial S}\right)_p \qquad \qquad V = \left(\frac{\partial H}{\partial p}\right)_S$$

- Also, in a flow process, if a **fluid** comes **in** at **pressure** p_1 and **volume** V_1 with **energy** U_1 and **leaves** at **pressure** p_2 and **volume** V_2 with **energy** U_2 , then it is **enthalpy** that is conserved.
- Helmholtz Free Energy F = U TS $\Rightarrow dH = -SdT pdV$
 - o The natural variables are now temperature and volume.
 - o For isothermal processes between V_1 and V_2 :

$$\Delta F = -\int_{V_1}^{V_2} p \,\mathrm{d}\, V$$

- **GIBBS FREE ENERGY** $\boxed{G = H TS}$ $\Rightarrow dG = -SdT + Vdp$
 - o The natural variables are now pressure and temperature.

- \circ If both T and p are constant for a process, G is conserved.
- Consider a system in contact with surroundings at T_o and p_o . Let's transfer energy dU and volume dV from surroundings \rightarrow system. Consider the internal energy of the surroundings, dU_o :

$$\begin{split} \mathrm{d}\, U_{\scriptscriptstyle 0} &= T_{\scriptscriptstyle 0}\,\mathrm{d}S_{\scriptscriptstyle 0} - p_{\scriptscriptstyle 0}\,(-\mathrm{d}\,V) = -\mathrm{d}\,U\\ \mathrm{d}S_{\scriptscriptstyle 0} &= -\frac{\mathrm{d}\,U + p_{\scriptscriptstyle 0}\mathrm{d}\,V}{T_{\scriptscriptstyle 0}} \end{split}$$

And now, let's assume that as a result of the change, the **entropy of** the system has changed by dS – then, assuming that the process does occur:

$$\begin{split} \mathrm{d}S_{\mathrm{tot}} &= \mathrm{d}S_{\mathrm{0}} \!+\! \mathrm{d}S \stackrel{2^{\mathrm{nd}} \, \mathrm{law}}{\geq} \widetilde{\mathbf{0}} \\ \Rightarrow \mathrm{d}\,U + p_{\mathrm{0}} \mathrm{d}\,V \!-\! T_{\mathrm{0}} \mathrm{d}S \leq \mathbf{0} \end{split}$$

If we define $A = U + p_0 V - T_0 S$, then this means that

 $\mathrm{d} A \leq 0$

So A always decreases for a process that occurs, at tends to a minimum at equilibrium.

It turns out that dA can be identified with the potentials above, whenever their natural variables are constant.

• The **Gibbs-Helmholtz Equations** can be derived by writing any one potential as the sum of others – for example, to obtain one for H in terms of F, write H = F + pV + TS, and convert the p and the S into forms involving F, using the differentials.

Maxwell Relations

- Derive them **directly** from the potentials.
- However, remember that
 - The **BOTTOM BITS** of the equation are the **NATURAL VARIABLES** of the potential they're from.
 - The **TOP BITS** are the **OTHER VARIABLES** of the potential they're from.

Strategy for these problems

• Write down a function of state in terms of particular variables. These can be "non-natural variables" by expressing one of the differentials as a differential itself... eg:

$$\begin{split} \mathrm{d}f &= \left(\frac{\partial f}{\partial x}\right)_z \,\mathrm{d}x + \left(\frac{\partial f}{\partial z}\right)_x \,\mathrm{d}z \\ &= \left(\frac{\partial f}{\partial x}\right)_z \,\mathrm{d}x + \left(\frac{\partial f}{\partial z}\right)_x \left[\left(\frac{\partial z}{\partial x}\right)_y \,\mathrm{d}x + \left(\frac{\partial z}{\partial y}\right)_x \,\mathrm{d}y\right] \\ &= \underbrace{\left[\left(\frac{\partial f}{\partial x}\right)_z + \left(\frac{\partial f}{\partial z}\right)_x \left(\frac{\partial z}{\partial x}\right)_y\right] \,\mathrm{d}x + \underbrace{\left(\frac{\partial f}{\partial z}\right)_x \left(\frac{\partial z}{\partial y}\right)_x \,\mathrm{d}y}_{(\partial f / \partial y)_x} \,\mathrm{d}y \end{split}$$

- Use Maxwell's relations to transform to easier partial differentials
- Use one of the following Theorems

$$\begin{pmatrix} \frac{\partial f}{\partial y} \\ \frac{\partial f}{\partial z} \\ \end{pmatrix}_x = \begin{pmatrix} \frac{\partial f}{\partial z} \\ \frac{\partial z}{\partial y} \\ \end{pmatrix}_x \text{ (chain rule}$$

$$\begin{pmatrix} \frac{\partial x}{\partial z} \\ \frac{\partial y}{\partial z} \\ \frac{\partial y}{\partial z} \\ \end{pmatrix}_y = \frac{1}{(\partial z / \partial x)_y} \text{ (reciprocal theorem)}$$

$$\begin{pmatrix} \frac{\partial x}{\partial y} \\ \frac{\partial z}{\partial z} \\ \frac{\partial z}{\partial z} \\ \frac{\partial z}{\partial y} \\ \frac{\partial z}{\partial z} \\ \frac{\partial z}{\partial y} \\ \frac{\partial$$

- Identify a heat capacity
- Identify a generalised susceptibility

This quantifies how much a particular variable changes when a generalised force (the differential of U w.r.t. some other parameter) is applied. For example, compressibilities (isothermal and adiabatic) are

$$\kappa_{\scriptscriptstyle T} = -\frac{1}{V} \! \left(\frac{\partial \, V}{\partial p} \right)_{\! T} \qquad \qquad \kappa_{\scriptscriptstyle S} = -\frac{1}{V} \! \left(\frac{\partial \, V}{\partial p} \right)_{\! S}$$

We can show that $\kappa_T / \kappa_S = \gamma > 1$, which implies that **adiabats** are steeper than isotherms.

Statistical Thermodynamics

Basic Ideas

- A multi-particle system can be defined as a set of quantum states.
- A particular configuration (quantum state) is called a **microstate**
 - Each microstate is **equally likely** to occur
 - The internal dynamics of the system are such that it can make **transitions** between microstates.
 - Given **enough time**, the system will **explore all** possible microstates and spend an **equal time** in **each of them** (the **ergodic hypothesis**).

Taken together, these three facts imply that a system will appear to choose a macroscopic configuration which maximizes the number of microstates.

- Measurements concern a **property** of the **macrostate** of the system, comprising of many different microstates, each with the same macroscopic property.
- The number of microstates making up a macrostates is given the symbol Ω .
- It is possible to find Ω as a function of some variable of the system around Ω₀, the maximum, and it is usually found that it peaks extremely sharply at that point the highest probability configuration gives the macroscopic state to a high degree of accuracy.

Temperature

• The 0^{th} Law of Thermodynamics stated that

Two systems, each separately in thermal equilibrium with a third, are also in thermal equilibrium with each other

This allows a definition of temperature as "the property which systems in thermal equilibrium have in common".

In a way this law states that "thermometers work", because it allows us to measure the temperature of one body by placing it in contact with another body which displays a property that has well-known dependence on temperature.

- Consider two systems in thermal contact with each other
 - o With energy (a property of the **macrostate**) E_1 and E_2
 - $\circ~$ With each of these energy macrostates corresponding to $\Omega(E)$ microstates.
 - $\circ \quad \text{With total energy } E = E_1 + E_2 \Rightarrow \mathrm{d} E_2 \, / \, \mathrm{d} E_1 = -1 \, .$
- The whole system can therefore be in any one of $\Omega(E_1)\Omega(E_2)$ microstates.
- If those two systems have come to Thermal Equilibrium, the number of microstates the whole system is in must be at a maximum. So, maximising the expression with respect to E₁:

$$\begin{aligned} \frac{\mathrm{d}}{\mathrm{d}E_1} \big[\Omega(E_1) \Omega(E_2) \big] &= 0\\ \frac{1}{\Omega_1} \frac{\mathrm{d}\Omega_1}{\mathrm{d}E_1} = \frac{1}{\Omega_2} \frac{\mathrm{d}\Omega_2}{\mathrm{d}E_2}\\ \frac{\mathrm{d}\ln\Omega_1}{\mathrm{d}E_1} &= \frac{\mathrm{d}\ln\Omega_2}{\mathrm{d}E_2} \end{aligned}$$

• So we see that the function $d \ln \Omega / dE$ must be a **constant** of the system, and must therefore be a function of **temperature**. We define:

1	$- d \ln \Omega$	1
$k_{\scriptscriptstyle B}T$	dE	

The Boltzmann Distribution – Canonical Ensembles

• Consider a system in equilibrium with a reservoir at temperature T.



And, consider that

- The **total energy** is E, and the energy of the system is ε , which leaves $E \varepsilon$ to the reservoir.
- The **temperature** of the reservoir is **constant**, as long as ε is small.
- The number of **microstates** of the **reservoir** corresponding to energy $E - \varepsilon$ is $\Omega_r(E - \varepsilon)$, and the number of microstates of the **system** corresponding to energy ε is $\Omega_s(\varepsilon)$.

This is known as the **canonical ensemble**.

- We can view a gas, for example, as a canonical example, where one molecule is the system, and the rest of the gas is the reservoir.
- The number of **microstates** the system can be in is therefore

$$\Omega_{total}$$
(system has energy ε) = $\Omega_r(E - \varepsilon)\Omega_s(\varepsilon)$

• But since ε is small, we can perform a Taylor Expansion:

$$\ln\left[\Omega_r(E-\varepsilon)\right] = \ln\left[\Omega(E)\right] - \frac{\mathrm{d}\ln\left[\Omega(E)\right]}{\mathrm{d}E}\varepsilon + \underbrace{\frac{\mathrm{d}^2\ln\left[\Omega(E)\right]}{\mathrm{d}E^2}\varepsilon^2}_{\propto \frac{\mathrm{d}E}{\mathrm{d}E}=0, \text{ because of}\atop \mathrm{assumptions}} + \cdots$$

And using our definition of temperature:

$$\ln \left[\Omega_r(E-\varepsilon)\right] = \ln \left[\Omega(E)\right] - \frac{\varepsilon}{k_B T}$$
$$\Omega_r(E-\varepsilon) = \Omega(E)e^{-\varepsilon/k_B T}$$

• Combining this with the above, we get

$$\Omega_{total} \propto \Omega_s(arepsilon) e^{-arepsilon/k_B T}$$

• Now, we said above that the **probability** of a given **macrostate** is proportional to the number of microstates that make it up. So, normalising:

$$P(\text{system in energy } \varepsilon) = \frac{\Omega_s(\varepsilon)e^{-\varepsilon/k_BT}}{\sum_{\forall \varepsilon}\Omega_s(\varepsilon)e^{-\varepsilon/k_BT}}$$

This is the **Boltzmann Distribution**.

• In a way, the reason why the probability falls as ε increases is because the degeneracy of the reservoir decreases exponentially with ε .

The Boltzmann Distribution – Other Derivation

• Consider N distinguishable particles and a set of energy levels ε_j , each with degeneracy g_j and population N_j . Consider further that the whole system has energy U. We can write

$$\begin{split} N &= \sum_{\forall j} N_j \\ U &= \sum_{\forall j} N_j \varepsilon_j \end{split}$$

• The number of ways of choosing N_i particles to be in the state ε_1 is (taking degeneracy into account)

$$\Omega_{1} = \underbrace{\frac{N!}{(N-N_{1})!}}_{\substack{\text{Number of ways} \\ \text{of choosing } N_{1} \\ \text{from } N \\ \text{degeneracy later}}} \underbrace{\frac{1}{N_{1}!}_{\substack{\text{Degeneracy} \\ \text{factor}}} \underbrace{g_{1}^{N_{1}}}_{\substack{\text{Degeneracy} \\ \text{factor}}} = \frac{N!}{N_{1}!(N-N_{1})!}g_{1}^{N_{1}}$$

The number of ways of choosing N_2 of the remaining ones is

$$\Omega_2 = \frac{(N - N_1)!}{N_2!(N - N_1 - N_2)!}g_2^{N_2}$$

Note that when multiplied together, some cancelling occurs. Therefore, the total number of configurations is

$$\boldsymbol{\Omega} = \prod_{\forall j} \boldsymbol{\Omega}_j = N! \prod_{\forall j} \frac{g_j^{N_j}}{N_j!}$$

Notes:

- We considered the particles to be **distinguishable** we considered that there was **more than one way** for N_i particle to exist in the state ε_i . More on this later.
- $\circ~$ We assumed that the degeneracy factor was g_i^N . This is true only if
- Now, we "simply" need to minimise ln Ω with respect to the N_i, subject to the two constraints. The steps are roughly as follows:
 - o Take the logarithm. Simplify logarithms using Stirling's Formula:

$$\ln N! = N \ln N - N$$

- \circ Add the Lagrange undetermined multipliers a and b using the two constraints defined above.
- Differentiate with respect to an arbitrary N_i , and set to 0 most components of each sum will come out of the wash, because

they'll involve other terms than N_i . (Remember to use the product rule for the double sum).

• The result will now be something like

$$N_i = g_i \exp(a + b\varepsilon_i)$$

• To find b, feed this expression back into the expression for $\ln \Omega$, and obtain something like

$$\ln \Omega = C - bU \Rightarrow \frac{\partial \ln \Omega}{\partial U} = -b$$

And note that from our statistical definition of temperature

$$\frac{1}{kT} = \frac{\partial \ln \Omega}{\partial U}$$

Which simply gives

$$b = -\frac{1}{kT}$$

• The rest is simple manipulation, and taking a constant out.

Fermi-Dirac Statistics

- In this case, we consider the statistics of indistinguishable particles like fermions, where no two particles can occupy the same quantum state.
- In such a case
 - The number of ways of getting N_i particles into the ith level is 1

 all particles are identical, and so there's only one way of getting a certain number in.
 - Within the i^{th} level, the number of particles we have **must** be \leq the **degeneracy** of the level (g_i) , because no two particles can be in the same state. So $N_i \leq g_i$.
 - So Ω_i is the number of ways of arranging the g_i levels amongst the N_i particles:

$$\Omega_i = \frac{g_i!}{N_i!(g_i - N_i)!}$$

• The total number of microstates is therefore

$$\Omega = \prod_{i} \frac{g_i!}{N_i!(g_i - N_i)!}$$

• Solve as above (the constant a in the exponential is usually written $a = \mu/kT$, where μ is the <u>chemical potential</u>). This gives:

$$N_i = \frac{g_i}{e^{(\varepsilon_i - \mu)/kT} + 1}$$

Notes:

- At low temperature, μ ≫ kT, states for ε_i < μ are essentially fully filled (ie: N_i ≈ g_i), whereas those for ε_i > μ are essentially empty. As T → 0, the distribution approaches a step function, with μ being the Fermi energy the lowest energy up to which the states must be filled to accommodate all the particles.
- At high temperature $\mu \ll kT$, the chemical potential effectively has no effect.

Bose-Einstein Statistics

- In this case, we consider the statistics of indistinguishable particles like fermions, where two particles <u>can</u> occupy the same quantum state.
- In such a case
 - Once again, the **number** of ways of getting N_i particles into the i^{th} level is 1.
 - The number of ways of getting the N_i particles in the g_i different states is equivalent to having $N_j + g_j 1$ symbols in a line, $g_j 1$ of each are "|", and asking "how many different ways can they be shuffled".
 - o This is given by

$$\Omega_{j} = \frac{1}{N_{j}!} \frac{(N_{j} + g_{j} - 1)!}{(g_{j} - 1)!}$$

• Once again, calculations follow and we obtain (making the assumption that $g_i \gg 1$, to remove a "-1" from the denominator):

$$N_i = \frac{g_i}{e^{(\varepsilon_i - \mu)/kT} - 1}$$

Notes:

- $\circ~$ At high temperature, $\mu \ll kT$, this behaves like the Fermi-Dirac distribution.
- As $T \to 0$, the population becomes very small, expect for ε_i just above μ this is **Bose-Einstein Condensation**.

The Equipartition Theorem

• Let the energy, E of a particular system be given by

$$E = \alpha x^2$$

Assuming that x can take <u>any</u> value, the mean energy is given by

$$\begin{split} \langle E \rangle &= \int_{-\infty}^{\infty} EP(x) \, \mathrm{d}x \\ &= \frac{\int_{-\infty}^{\infty} \alpha x^2 e^{-\beta \alpha x^2} \, \mathrm{d}x}{\int_{-\infty}^{\infty} e^{-\beta \alpha x^2} \, \mathrm{d}x} \\ &= \frac{1}{2\beta} \\ &= \frac{1}{2} k_B T \end{split}$$

This result (the whole proof's a bit longer!) implies that each quadratic energy dependence of the system (mode of the system) has an energy $\frac{1}{2}k_BT$ associated with it. This is the Equipartition Theorem.

- There are, however, two assumptions involved in the derivation of this result:
 - That the wells are quadratic (the Harmonic Approximation).
 - That the parameter x can take **any** value, and that the variables could be **integrated continuously**. This is **dubious** for **quantum** systems with **quantised** values.

The Equipartition Theorem is only valid at <u>high</u> <u>temperatures</u> so that the <u>thermal energy</u> is <u>larger than</u> the <u>energy gap</u> between <u>quantised</u> <u>energy levels</u>.

The Partition Function

• The partition function is the sum over all states of the Boltzmann Factor:

$$Z = \sum_{\forall \alpha} e^{-\beta E_{\alpha}}$$

However, since the point from which we measure energy is arbitrary, Z is defined up to an arbitrary multiplicative constant.

- The partition functions only refers to a single particle in the system. It may well be coupled to a reservoir of other particles, but it's a single one.
- Z contains, "zipped up within it", full information about every property of the system
 - The internal energy

The **internal energy** is given by

$$U = \frac{\sum_{i} E_{i} e^{-\beta E_{i}}}{\sum_{i} e^{-\beta E_{i}}} = \frac{-\mathrm{d}Z / \mathrm{d}\beta}{Z} = -\frac{\mathrm{d}\ln Z}{\mathrm{d}\beta}$$

o The Entropy

We know that $P_j = e^{-\beta E_j} / Z$, and so using **Gibbs' Expression** for Entropy, we have

$$S = -k \sum_{i} P_{i} \ln P_{i}$$
$$= k \sum_{i} P_{i} (\beta E_{i} + \ln Z)$$
$$= k (\beta \langle E \rangle + \ln Z)$$
$$= k (\beta U + \ln Z)$$
$$S = \frac{U}{T} + k \ln Z$$

• The Helmholtz Energy

We know that
$$F = U - TS$$
, so

$$\boxed{F = -kT \ln Z}$$

$$Z = e^{-\beta F}$$

 \circ $\;$ The rest can then easily be worked out as follows...

$$S = -\left(\frac{\partial F}{\partial T}\right)_{V} = \frac{U - F}{T}$$
$$p = -\left(\frac{\partial F}{\partial V}\right)_{T}$$
$$H = U + pV$$
$$G = F + pV = H - TS$$
$$C_{V} = \left(\frac{\partial U}{\partial T}\right)_{V} = T\left(\frac{\partial S}{\partial T}\right)_{V}$$

• If the energy of a particular system depends on various independent contributions, $E_{i,j} = E_i^{(a)} + E_j^{(b)}$, then the combined Partition Function is given by

$$Z = \sum_{i} \sum_{j} e^{-\beta(E_{i}^{a} + E_{j}^{b})} = \sum_{i} \sum_{j} e^{-\beta E_{i}^{a}} e^{-\beta E_{j}^{b}} = \sum_{i} e^{-\beta E_{i}^{a}} \sum_{j} e^{-\beta E_{j}^{b}} = Z_{a} Z_{b}$$

So partition functions can be combined by multiplication.

The Ideal Monoatomic Gas

- Consider a gas in a cubic box of side *L*.
 - The wavefunction of each atom must be

$$\psi(x) = A\sin(k_x x)\sin(k_y y)\sin(k_z z)$$

Where

$$k_i = \frac{n_i \pi}{L}$$

And the n_i are **integers** (due to the boundary conditions).

 $\circ~$ In k-space, each state (distinct combination of $n {\rm s})$ occupies a volume

$$\frac{\pi^3}{L^3}$$

• The number of states with k in the range k + dk is therefore equal to the number of states in the relevant **octant** in k-space

$$g(k) dk = \frac{1}{8} \times 4\pi k^2 dk \times \frac{1}{\pi^3 / L^3} = \frac{Vk^2}{2\pi^2} dk$$

• If *m* is the **mass** of an atom, then the **kinetic energy** of the atom is given by

$$E = rac{\hbar^2 \left| m{k}
ight|^2}{2m} = rac{\hbar^2 \pi^2}{2m L^2} \left(n_x^2 + n_y^2 + n_z^2
ight)$$

• The partition function for a single molecule of our ideal gas is therefore

$$\begin{split} \sum_{i} g(E_i) e^{-\beta E_i} &= \int_0^\infty e^{-\beta E(k)} g(k) \,\mathrm{d}k \\ &= \int_0^\infty e^{-\beta \hbar^2 k^2 / 2m} \frac{V k^2}{2\pi^2} \,\mathrm{d}k \\ &= \frac{V}{\hbar^3} \left(\frac{m k_B T}{2\pi}\right)^{3/2} \end{split}$$

For a gas of **more than 1 indistinguishable molecule**, the partition function is

$$Z_N = \frac{Z^N}{N!}$$

This is because we can **combine** the **energies** as we did in the previous section, but we must make sure we do not **overcount states**. If we assume that the gas is **dilute** and that **no more than 1 particle occupies each state**, then we have **overcounted** by **exactly** *N*! This explains the denominator above.

- We can derive a number of properties for the partition function, all of which turn out to be as expected classically.
- We can also work out an expression for **entropy**

 $S = \frac{3}{2}Nk_B \ln T + Nk_B \ln \left(\frac{V}{N}\right) + \frac{3}{2}Nk_B \ln \left(\frac{mk_B}{2\pi\hbar^2}\right) + \frac{5}{2}Nk_B$

This is called the **Sackur-Tetrode Equation**, and it could not have been obtained using classical thermodynamics. It can also be used to demonstrate the **Gibbs Paradox**:

- The Sackur-Tetrode Equation correctly predicts the entropy change of a Joule expansion to be $\Delta S = Nk_B \ln 2$.
- However, if a vessel is divided into two, with an identical gas on each side, the entropy change when the partition is removed becomes $\Delta S = 0$.
- However, one could argue that the gas on **each side** of the partition has itself undergone a **Joule expansion**!
- This demonstrates the importance of realising that indistinguishable particles really are indistinguishable – no information can be lost when removing the partition in that example.
- Adiabatic changes
 - Consider reversibly compressing a gas, keeping the relative populations of the energy levels fixed ie: adiabatically compress it.
 - The energies are proportional to $V^{-2/3}$, so they'll increase, and $U \propto V^{-2/3}$.
 - However, $U = \frac{3}{2}kT$.
 - $\circ~$ Therefore, $TV^{3/2}={\rm constant}$, as expected for an adiabatic expansion.

• Heat and work – we can also consider a general change in the internal energy

$$\mathrm{d}\, U = N \sum p_i \mathrm{d} \varepsilon_i + \varepsilon_i \mathrm{d} p_i$$

This includes two terms

- The first involves changing energy levels but keeping populations fixed – this is an adiabatic change, and can be interpreted as the reversible work dW = -p dV.
- The second involves keeping the energy levels fixed but changing the populations and therefore the entropy. This can be interpreted as the reversible heat dQ = T dS.

The Ideal Diatomic Gas

• If the gas is now **diatomic**, we have two additional terms in the energy

$$E = \frac{\hbar^2 \pi^2}{2mL^2} \left(n_x^2 + n_y^2 + n_z^2 \right) + \frac{\ell(\ell+1)\hbar^2}{2I} + \underbrace{(n+\frac{1}{2})\hbar\omega}_{\text{vibrational E}} + \underbrace{(n+\frac{1}{2})\hbar\omega}_{\text{vibrational E}} \right)$$

- We need to take **products** to find the combined partition functions (see above) and the new contributions to the potentials are therefore **additive**.
- The actual contributions can be found by writing the **expressions** for the **rotational** and **vibrational partition funtions**. Notes:
 - The degeneracy of each rotational energy is $2\ell + 1$.
 - The **degeneracy** of each **vibrational energy** is 1.
- In each case, the limits $T \to 0, \infty$ need to be considered, because the general case is rather difficult to evaluate.

Paramagnetism

- Atoms in **crystals** may have **permanent magnetic dipole moments**. These are linked to their **angular momentum** (orbital or spin).
- In QM, angular momentum is quantised along any given axis as $J_z = m\hbar$ where *m* takes integer spaced values from *J* to -J, where *J* is the total angular momentum quantum number of the atom.

• The upshot of this is that the *z*-component of magnetic moment is also quantised, according to

$$\mu_z = g\mu_B m$$

Where $\mu_B = e\hbar/2m_e$ is the **Bohr Magneton** and g is the "Landé g factor", which is O(1).

- The energy of such a dipole in a magnetic field B along the z-axis is $\mu_z B$.
- The calculations reveal a number of interesting things (note that μ is the magnitude of the dipole moment)
 - o Curie's Law as $T \to \infty$, $\langle \mu_z \rangle \to \mu^2 B / kT$.
 - As $T \to 0$, $\langle \mu_z \rangle \to \mu$ (ie: all magnetic dipoles aligned with the field).
- We can obtain the magnetisation, M and magnetic susceptibility, χ of the magnet, given by

$$\begin{split} M &= n \left< \mu_z \right> \\ \chi &= \frac{M}{H} \approx \frac{\mu_0 M}{B} \end{split}$$

Photons, etc...

Introduction

- Consider an evacuated box of volume V at temperature T a cavity. It contains electromagnetic radiation.
- In thermodynamics, we can treat this as a gas of photons which is constantly being emitted/absorbed/reflected at the walls.

Spectral Energy Density

• It is relatively simple to work out that if the **photon gas** contains *n* **photons per unit volume**, the **total energy per unit volume** in the cavity is

$$u = n \langle \hbar \omega \rangle$$

However, one might ask which frequency ranges this energy is stored in.
 We therefore introduce the spectral energy density, u_λ, such that the energy stored in wavelengths of range λ to λ + dλ is

$$u_{\lambda} \,\mathrm{d}\lambda$$

The **total** energy density is then

$$u = \int u_{\lambda} \, \mathrm{d}\lambda$$

We can show that this spectrum is totally independent of the kind of cavity (shape/size/material) and is a universal function of λ and T by considering two cavities, each kept at T by a heat bath, and connected by a tube containing a filter that only allows wavelength in the range [λ, λ + dλ] though:



By the second law, no heat can flow along the tube (equal temperatures). Therefore, any energy flux from right to left must be balanced by a flux from left to right. Bearing in mind our expression for P (power absorbed per unit area) below, we find that $P_{left} = P_{right}$, and therefore that $u_{l,left} = u_{l,right}$.

Banalities

First, we note that if the photons are moving in completely arbitrary directions, then the fraction of them moving at angles θ to θ + dθ to any particular direction is

$$\frac{\mathrm{d}\Omega}{4\pi} = \underbrace{\frac{1}{4\pi}}_{\substack{\mathrm{Total \ solid}\\\mathrm{angles}}} \underbrace{2\pi\sin\theta\,\mathrm{d}\theta}_{\substack{\mathrm{Solid\ angle\ d\Omega\ subtented}\\\mathrm{solid\ directions}}} = \frac{1}{2}\sin\theta\,\mathrm{d}\theta$$

Therefore, the number of molecules per unit volume that are:

- Travelling with a speed v to v + dv.
- Travelling at angles θ to $\theta + d\theta$ to a certain direction. Is given by

$$\frac{1}{2}nf(v)\sin\theta \,\mathrm{d}v\,\mathrm{d}\theta = \frac{1}{2}n\sin\theta\,\mathrm{d}\theta$$
In the photon case,
where $v=c$

Now, if we consider this "**certain direction**" to be perpendicular to a wall, and if we consider particles:

- o Travelling with a speed v.
- Travelling at an angle θ to the wall.

Then only those particles **closer than** $v\cos\theta dt$ to the wall will **hit** it in **time** dt. In other words, those particles in a volume $v\cos\theta dt$. The **number** of such particles is therefore (from above)

$$\underbrace{v \cos \theta \, \mathrm{d}t}_{\text{Volume}} \underbrace{\frac{1}{2} nf(v) \sin \theta \, \mathrm{d}v \, \mathrm{d}\theta}_{\text{Particles per unit volume}} = \frac{1}{2} nv f(v) \cos \theta \sin \theta \, \mathrm{d}v \, \mathrm{d}\theta \, \mathrm{d}t$$
$$= \underbrace{\frac{1}{2} nc \cos \theta \sin \theta \, \mathrm{d}\theta}_{\frac{1}{2} nc \cos \theta \sin \theta \, \mathrm{d}\theta}$$

To find, therefore, the **total flux of photons** hitting **unit surface area** in **unit time**, we integrate this over **all angles** [Note: in this context, "all angles" means $0 \rightarrow \pi/2$, because of the way we defined solid angles above – the range $-\pi/2 \rightarrow 0$ is implicitly included]:

$$\Phi = \frac{1}{2} \int_0^{\pi/2} nc \cos \theta \sin \theta \, \mathrm{d}\theta$$
$$\Phi = \frac{1}{4} nc$$

Since each **photon** has energy $\langle \hbar \omega \rangle$, we can also write the **average power incident per unit area of wall** as

$$P = \hbar\omega\Phi = \frac{1}{4}uc$$

(See next section for definition of u).

• We saw above that the **number of photons** travelling at angles $[\theta, \theta + d\theta]$, in a photon gas, hitting **unit area** of a wall in **unit time** was $\frac{1}{2}nc\cos\theta\sin\theta\,d\theta$

Now, if instead of the number of photons, n, we are interested in the "momentum change perpendicular to the wall at each collision and relfection", then we must replace n

$$n \to \int_{\forall \varepsilon} \underbrace{2}_{\substack{\text{Double the momentum, because we're recoiling \\ \text{momentum is because we're recoiling }}} \underbrace{\frac{u_{\varepsilon}}{c}}_{\substack{\text{Momentum is } \\ \text{energy/c} }} \underbrace{\cos \theta}_{\substack{\text{Perpendicular Integrate over } \\ \text{to wall all energies }}} \underbrace{\mathrm{d}\varepsilon}_{\text{all energies }} = \frac{2u\cos \theta}{c}$$

And we now get that the "momentum change per unit area per unit time = pressure" is

$$p = \int_0^{\pi/2} \frac{2u\cos\theta}{c} \frac{1}{2}c\cos\theta\sin\theta\,\mathrm{d}\theta$$
$$= \int_0^{\pi/2} u\cos^2\theta\sin\theta\,\mathrm{d}\theta$$
$$\boxed{p = \frac{1}{3}u}$$

The Stefan-Boltzmann Law

• Consider the **First Law of Thermodynamics** - dU = TdS - pdV. Differentiating w.r.t V at constant T:

$$\left(\frac{\partial U}{\partial V}\right)_{T} = T \left(\frac{\partial S}{\partial V}\right)_{T} - p = T \left(\frac{\partial P}{\partial T}\right)_{V} - p$$

However, $(\partial U / \partial V)_T = u$ [by definition, but this can also be proved by noting that since U = uV, $(\partial U / \partial V)_T = u + V(\partial u / \partial V)_T = u$], so:

$$u = T \left(\frac{\partial p}{\partial T} \right)_V - p$$

But we have shown that $p = \frac{1}{3}u$, so

$$u = \frac{1}{3} \left[T \frac{\mathrm{d}u}{\mathrm{d}T} - u \right] \Rightarrow \frac{\mathrm{d}u}{u} = 4 \frac{\mathrm{d}T}{T}$$

$$u = A T^4$$

Feeding this into our expression for **power incident per unit area**:

$$P = \frac{1}{4}uc = \left(\frac{1}{4}Ac\right)T^4 = \sigma T^4$$

• Note, also, that if the **cavity** is in **equilibrium**, "**power incident** = **power emitted**", so the above also gives an expression for the power emitted.

Kirchhoff's Law

- We wish to discuss how well particular surfaces of a cavity will absorb or emit electromagnetic radiation of a particular wavelength. We therefore define:
 - The spectral absorptivity α_{λ} is the fraction of incident radiation absorbed at λ .
 - The spectral emissive power e_{λ} is a function such that the power emitted per unit area of the surface having wavelength in the range $[\lambda, \lambda + d\lambda]$ is

$$e_{\lambda} \,\mathrm{d}\lambda$$

Now, in equilibrium, the power absorbed is equal to the power emitted, so using our expression for the power incident above, we have that in the range [λ, λ + dλ]:

$$\begin{pmatrix} \frac{1}{4} u_{\lambda} \, \mathrm{d}\lambda \, c \\ \alpha_{\lambda} &= e_{\lambda} \, \mathrm{d}\lambda \\ \hline \frac{e_{\lambda}}{\alpha_{\lambda}} &= \frac{c}{4} \, u_{\lambda} \end{bmatrix}$$

This is **Kirchhoff's Law** – it states that $e_{\lambda} / \alpha_{\lambda}$ is a **universal function** of λ and T. Therefore, if we fix T, then $e_{\lambda} \propto \alpha_{\lambda}$. This means that "good absorbers are good emitters" and vice versa.

- For a **perfect black body**, $\alpha_{\lambda} = 1$ for all λ , which also means that the black body is the **best possible emitters**.
- A cavity whose walls have $\alpha_{\lambda} = 1$ is known as a black body cavity.

Statistical Mechanics

- The **black body spectrum** can be derived in the usual ways (see QM course).
- To **convert** the spectrum from one kind of measurement (eg: **wavelength**) to another (eg: **frequency**), write (for example)

$$u_{\lambda} \, \mathrm{d}\lambda = u_{\nu} \, \mathrm{d}\nu$$

• This can be **integrated** to give

$$u = \frac{\hbar}{\pi^2 c^3} \int_0^\infty \frac{\omega^3}{e^{\beta\hbar\omega} - 1} \mathrm{d}\omega = \frac{k^4 \pi^2}{15\hbar^3 c^3} T^4$$

The Stefan-Boltzmann Law. We have therefore obtained

$$\sigma = \frac{\pi^2 k_B^4}{60c^2 \hbar^3}$$